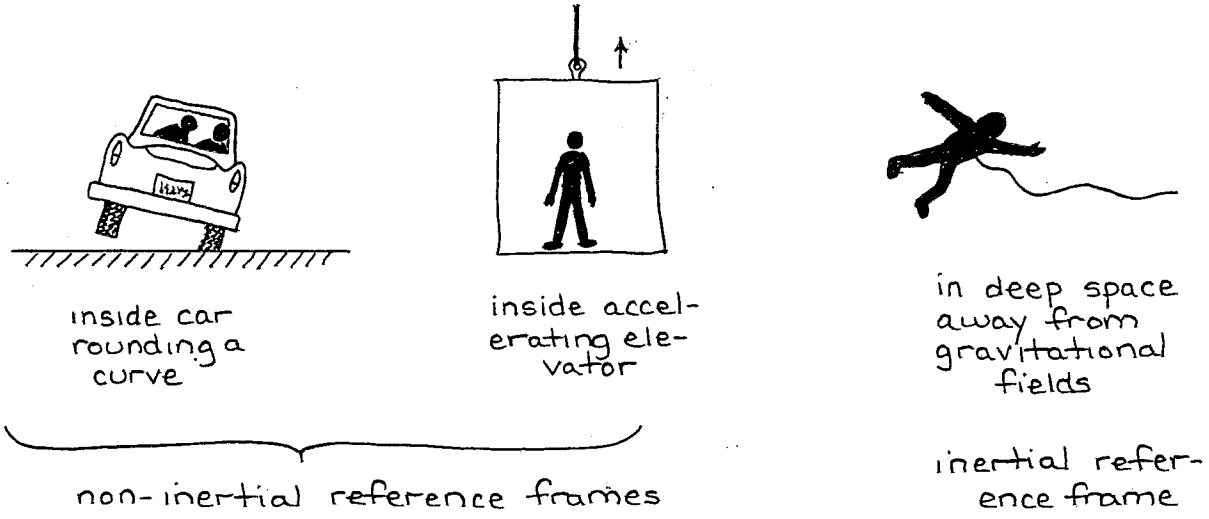


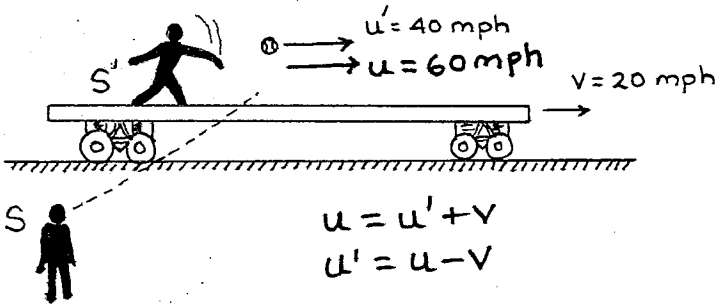
RELATIVITY

inertial frame of reference - any frame of reference in which Newton's first law of motion holds (isolated object does not accelerate).



principle of classical relativity - the laws of (Newtonian) mechanics are the same in all inertial reference frames

addition of velocities (classical)



2 inertial frames of reference

S = frame of reference of man standing at station

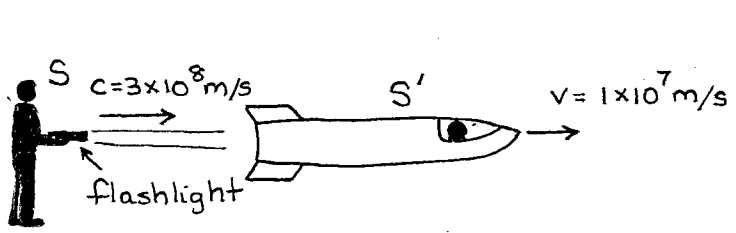
S' = frame of reference of man on flatcar

u' = velocity of baseball as measured by thrower in S' ($u' = 40 \text{ mph}$)

u = velocity of baseball as measured by man at station in S ($u' = u - v$)

$$u = u' + v = 40 \text{ mph} + 20 \text{ mph} = 60 \text{ mph}$$

or

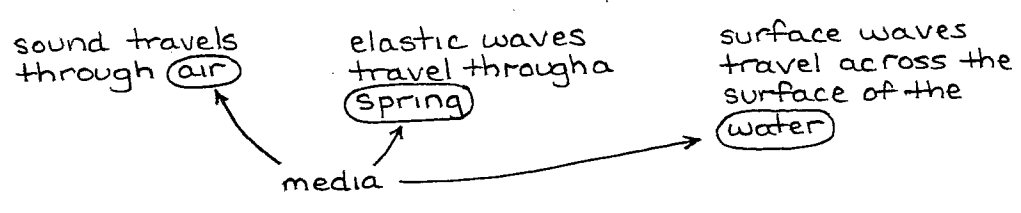


$$c' = c - v = 2.9 \times 10^8 \text{ m/s}$$

* classical relativity and addition of velocities "makes sense", but it is not totally correct

Maxwell's Equations predict a speed of 3.0×10^8 m/s for light in a vacuum -- but relative to what reference frame?

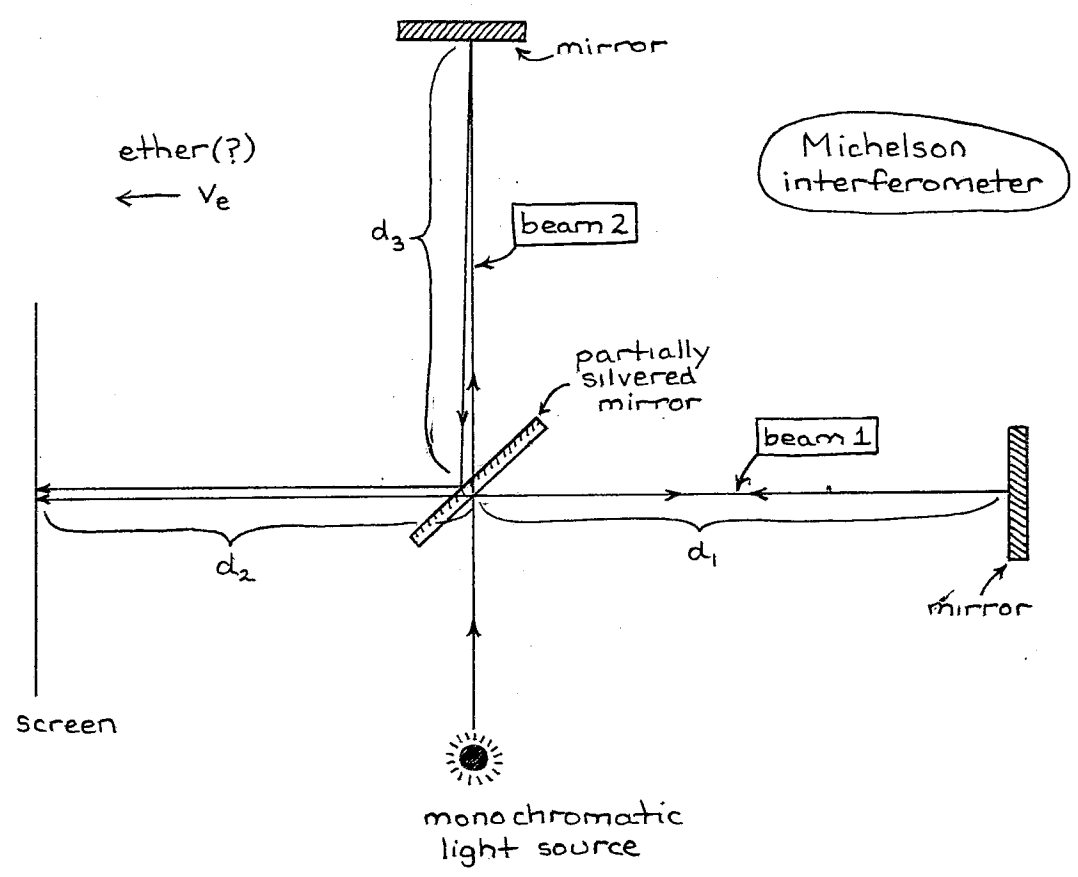
* any speed must be referenced to some reference frame in order for it to have any meaning



* from 1870-1905 many physicists believed a medium was necessary to conduct electromagnetic radiation; this was called ether or luminiferous ether

Michelson-Morley Experiment (devised by Albert Michelson, American)

* to detect and measure the speed of the ether



$t_1 =$ time for beam 1 to reach screen

$$t_1 = \frac{d_1}{c-v_e} + \frac{d_1}{c+v_e} + \frac{d_2}{c+v_e}$$

$t_2 =$ time for beam 2 to reach screen

$$t_2 = \frac{d_3}{c} + \frac{d_3}{c} + \frac{d_2}{c+v_e}$$

time of arrival difference $\Delta t = t_1 - t_2$

$$\Delta t = \frac{d_1}{c-v_e} + \frac{d_1}{c+v_e} - \frac{d_3}{c} - \frac{d_3}{c} = \frac{d_1 c}{c^2 - v_e^2} - \frac{2d_3}{c}$$

rotate interferometer
90° clockwise

$$t'_1 = \frac{d_1}{c} + \frac{d_1}{c} + \frac{d_2}{c}$$

$$t'_2 = \frac{d_3}{c-v_e} + \frac{d_3}{c+v_e} + \frac{d_2}{c}$$

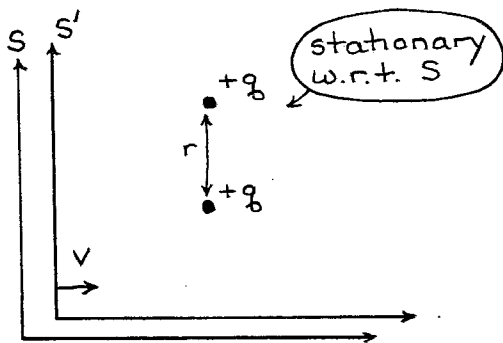
$$\Delta t' = t'_1 - t'_2 = \frac{2d_1}{c} - \frac{d_3 c}{c^2 - v_e^2}$$

* Since Δt and $\Delta t'$ will be different in general, different interference patterns will be seen for each of the configurations (rotations).

* The shift in fringe patterns observed would allow the computation of v_e

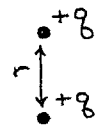
* Result of Michelson-Morley experiment: no fringe shift!!

apparent violation of classical relativity



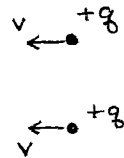
what S sees:

$$F_{rep} = \frac{kq^2}{r^2}$$



what S' sees:

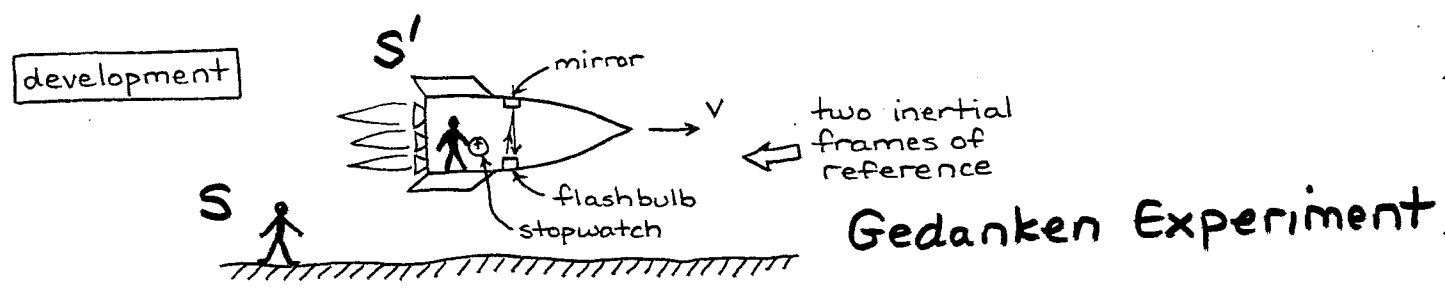
$$F_{rep} = \text{less than } \frac{kq^2}{r^2}$$



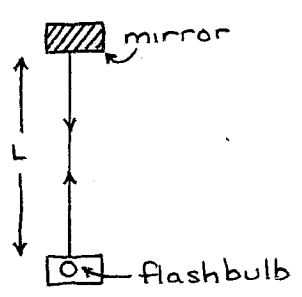
(magnetic force added)

special theory of relativity (A. Einstein, 1905) STR.

- (1) principle of relativity (postulate): the laws of physics are the same in all inertial frames of reference
- (2) constancy of the speed of light (postulate): the speed of light in a vacuum is constant w.r.t. inertial frames of reference



what the spaceman sees

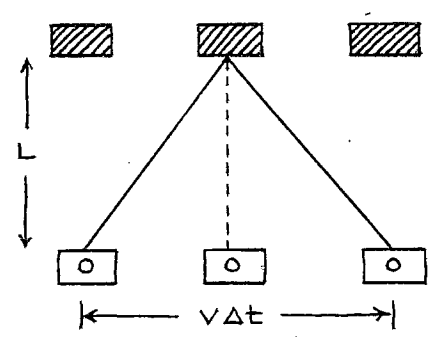


time between flash and light pulse's return to point of origin as measured by spaceman

$$\Delta t_0 = \frac{\text{distance}}{\text{speed}} = \frac{2L}{c}$$

$$\Delta t_0 = \frac{2L}{c}$$

what the earthling sees



$$\Delta t = \frac{\text{distance}}{\text{speed}} = \frac{2\sqrt{L^2 + (\frac{v\Delta t}{2})^2}}{c}$$

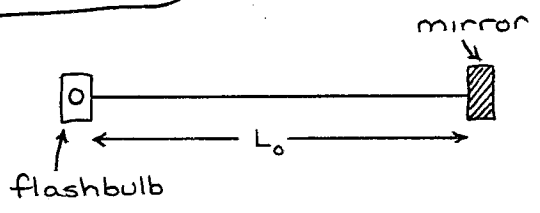
$$c^2\Delta t^2 = 4(L^2 + \frac{v^2\Delta t^2}{4})$$

$$c^2\Delta t^2 - v^2\Delta t^2 = 4L^2$$

but $L = \frac{c\Delta t_0}{2} \rightarrow c^2\Delta t^2 - v^2\Delta t^2 = c^2\Delta t_0^2$

time dilation $\rightarrow \Delta t = \frac{\Delta t_0}{\sqrt{1 - v^2/c^2}}$

what the spaceman sees

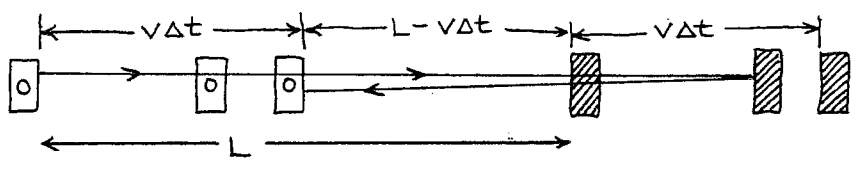


time between flash and return of pulse to point of origin as seen by spaceman

$$\Delta t_0 = \frac{\text{distance}}{\text{speed}} = \frac{2L_0}{c}$$

$$\Delta t_0 = \frac{2L_0}{c}$$

what the earthling sees



$$\Delta t = \frac{L}{c-v} + \frac{L}{c+v}$$

$$= \frac{2cL}{c^2 - v^2} = \frac{2L\gamma^2}{c}$$

$$\Delta t = \frac{2L\gamma^2}{c}$$

* using time dilation ($\Delta t = \Delta t_0 \gamma$)

$$\Delta t = \frac{2L\gamma^2}{c} = \Delta t_0 \gamma = \frac{2L_0}{c} \gamma \Rightarrow$$

$$L = \frac{L_0}{\gamma}$$

Lorentz-FitzGerald contraction

eg an observer sees a spaceship that is 100m long when at rest; if this ship were passing the observer at 0.5c, how long will the observer measure it to be?

relativistic correction to momentum

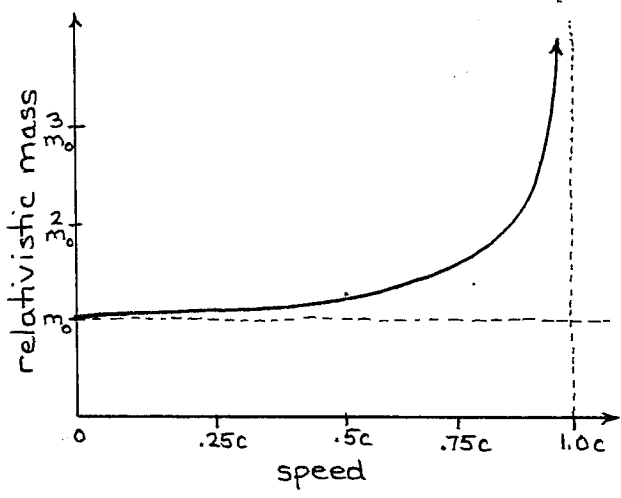
$p = m_0 v$
↙ classical

$p = m_0 \gamma v$
↙ relativistic

conservation of momentum - in any isolated system the total relativistic momentum remains constant

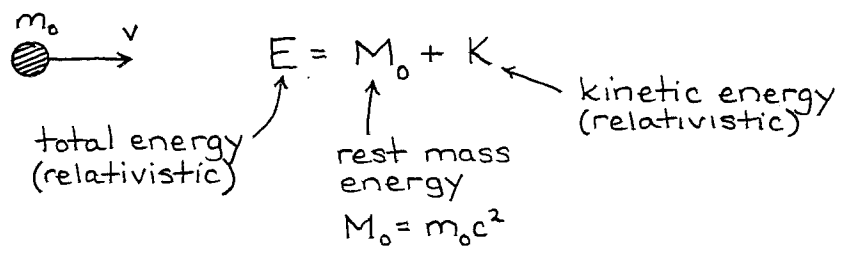
$p_{rel} = \overbrace{m_0 \gamma}^{\text{relativistic mass}} v$ ← often denoted m
↙ rest mass

relativistic mass = $m_0 \gamma = \frac{m_0}{\sqrt{1 - v^2/c^2}} = m$



energy and momentum

* the total energy (relativistic) of a mass can in one sense be thought of as composed of 2 components



$E = mc^2 = m_0 \gamma c^2$
 $M_0 = m_0 c^2$
 $K = E - M_0 = mc^2 - m_0 c^2$
 $= m_0 (\gamma - 1) c^2$

eg Compute the rest mass energy of a proton

$$M_p = m_p c^2 = (1.67 \times 10^{-27} \text{ kg}) (3 \times 10^8 \text{ m/s})^2 = \boxed{1.5 \times 10^{-10} \text{ J}}$$

$$= \boxed{938 \text{ MeV}}$$

* In high energy physics, MeV (million electron volts) is a commonly used energy and mass-energy unit

- $M_{\text{electron}} = .511 \text{ MeV}$
- $M_{\text{proton}} = 938.3 \text{ MeV}$
- $M_{\text{neutron}} = 939.6 \text{ MeV}$

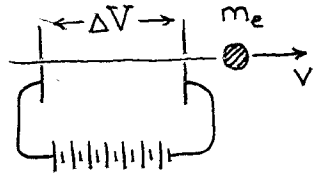
eg How much energy is required to accelerate an electron from rest to .95c?

$$K = M_e (\gamma - 1) c^2 = M_e (\gamma - 1) = (.511 \text{ MeV}) \left(\frac{1}{\sqrt{1 - .95^2}} - 1 \right) = \boxed{1.13 \text{ MeV}}$$

What would the classical kinetic energy be?

$$K_{\text{classical}} = \frac{1}{2} m_e v^2 = \frac{1}{2} (m_e c^2) \left(\frac{v}{c} \right)^2 = \frac{1}{2} (.511 \text{ MeV}) (.95)^2 = .23 \text{ MeV}$$

eg An electron is accelerated across a 2,000,000 V potential difference. What is its final speed?



$$K = q \Delta V = M_e (\gamma - 1)$$

$$\gamma - 1 = \frac{q \Delta V}{M_e}$$

$$\gamma = \frac{q \Delta V}{M_e} + 1 = \frac{2 \times 10^6 \text{ eV}}{.511 \times 10^6 \text{ eV}} + 1 = 4.9139$$

$$\frac{1}{\sqrt{1 - v^2/c^2}} = 4.9139$$

$$1 - v^2/c^2 = \frac{1}{4.9139^2}$$

$$v^2 = \left(1 - \frac{1}{4.9139^2} \right) c^2 = 8.627 \times 10^{16}$$

$$\boxed{v = 2.94 \times 10^8 \text{ m/s} = .979 c}$$

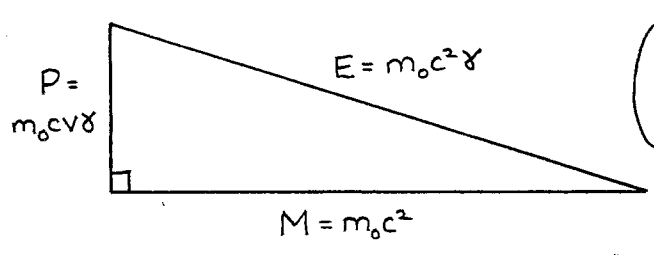
note: relativistic momentum is $p = m_0 v \gamma$

define (relativistic) momentum-energy as $P = pc = m_0 v c \gamma$

then $P^2 = m_0^2 c^2 v^2 \gamma^2 = M^2 \frac{v^2}{c^2} \gamma^2$

and $M^2 + P^2 = M^2 + M^2 \frac{v^2}{c^2} \gamma^2 = M^2 \underbrace{\left(1 + \frac{v^2}{c^2} \gamma^2\right)}_{\gamma^2} = M^2 \gamma^2 = (M \gamma)^2 = E^2$
total energy \uparrow

so $E^2 = M^2 + P^2$ or $E^2 = (m_0 c^2)^2 + (m_0 v \gamma c)^2$



mass energy and momentum energy "add at right angles" to give total energy

general theory of relativity (A. Einstein, 1915)

* special relativity extended for use in acceleration (esp. gravitational) fields

principle of equivalence - no experiment performed in a closed, accelerating system can distinguish between the effects of a uniform gravitational field and the effects of a uniform acceleration

implications of GTR